

Fermilab

Transverse Beam Heating Distributions

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## Transverse Beam Heating Distributions

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A beam is heated in one dimension with a dipole kicker which deflects the beam by an angle  $\theta(t)$ , at a location where the lattice functions are  $(\alpha,\beta,\gamma)$ .  $\theta$  has a power spectrum  $P(\omega)$  in the sense that  $\theta^2$  represents power. We will use variables L, $\gamma$  defined as

$$I = (\gamma x^2 + 2\alpha x x' + \beta x'^2)/2$$

$$tan \gamma = \alpha + \beta x'/x$$
1.

When a particle is deflected by an angle  $\theta$ , I and  $\gamma$  change by

$$\delta I = (2I\beta)^{1/2} \sin \theta + \beta \theta^2 / 2$$

$$\delta Y = (2\beta/I)^{1/2} \cos \theta \qquad 2.$$

Although  $\gamma$  will diffuse, we are only interested in amplitudes I. To use the Fokker-Planck (FP) equation, we need to calculate the sum over many revolutions of the 6I's (= $\Delta$ I) for a particle, and find those parts of  $\Delta$ I and  $(\Delta I)^2$  which are proportional to time. The particle, which has (angular) revolution frequency  $\omega_0$  and betatron frequency  $\omega_\beta = \omega_0 \nu$  passes the kicker at times  $t_Q$  with betatron phase  $\gamma_0$  where

$$t_{\varrho} = t_{o} + 2\pi \ell/\omega_{o}$$

$$\gamma_{\varrho} = \gamma_{o} + \omega_{e}t_{\varrho}$$
3.

The final values of  $\Delta I$  and  $(\Delta I)^2$  are then averaged over  $t_o$  and  $\gamma_o$ . We find;

$$\begin{split} \Delta I/T &= \beta \omega_0 \int P(\omega) d\omega/2\pi \equiv k_w \\ (\Delta I)^2/T &= \beta I \omega_0^2 \sum_{p=0}^{\infty} P(|p\omega_0 \pm \omega_g|)/2\pi \equiv 2kI \end{split} \qquad 4. \end{split}$$

The FP equation is

$$\frac{\partial f}{\partial t} + \frac{\partial \Phi}{\partial l} = 0$$

$$\Phi = k_w f - k \frac{\partial f}{\partial l} = 0$$
5.

For narrow band heating, the first term can be neglected, since its contribution to FP is of order  $BW/f_0$  compared to that of the second term. For wideband heating  $BW>f_0$ , it changes the nature of the diffusion, and the subsequent distribution. Eq. 5 admits solutions, where C is a constant,

$$f = Ce^{-\alpha t} J_0(2\sqrt{(\alpha I/k)})$$
 6.

Now f must be 0 at, as well as beyond, the aperture limit  $I_m$ ; otherwise its infinite derivative there would imply infinite flux. Then  $J_0(2\sqrt{(\alpha I_m/k)}) = 0$ , and the allowed values of  $\alpha$  are

$$\alpha_{i} = k j_{i}^{2} / 4 l_{m}$$
 7.

Here  $j_i$  is the  $i\underline{th}$  zero of  $J_0$ . Since the set of functions  $J_0(j_ix)$  is complete in the interval (0,1), the initial distribution can be expressed as a series of  $J_0(j_i\sqrt{(I/I_m)})$ . Each term then decays with its own lapse rate  $\alpha_i$ . For our purposes, we note that all terms above i=1 decay very rapidly compared to i=1, so the distribution quickly approaches  $J_0(j_1\sqrt{(I/I_m)})$ .

What one measures with the scraper is the integral distribution F(y) as a function of y, where  $y=\sqrt{(1\beta_S)}$ , and  $\beta_S$  is the value at the scraper. F(y) is given by

$$F(y) = yJ_1(j_1y/y_m)/[y_mJ_1(j_1)]$$
 8.

This can be compared directly to the plots on the Lexidata.